

Gravity Probe B Overview

Barry Muhlfelder

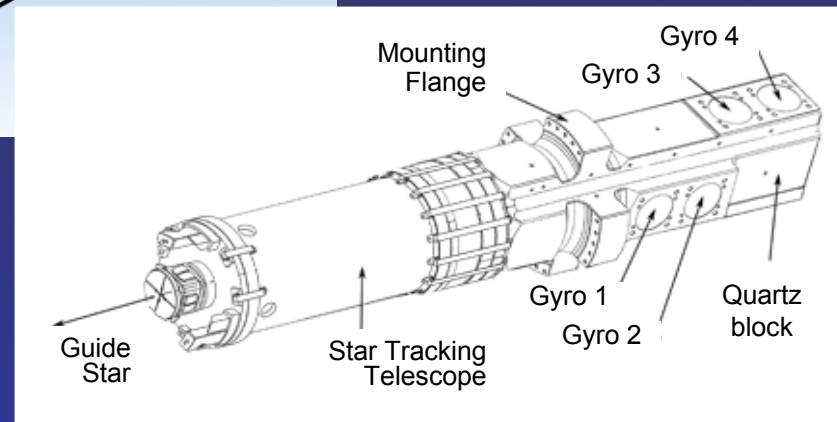
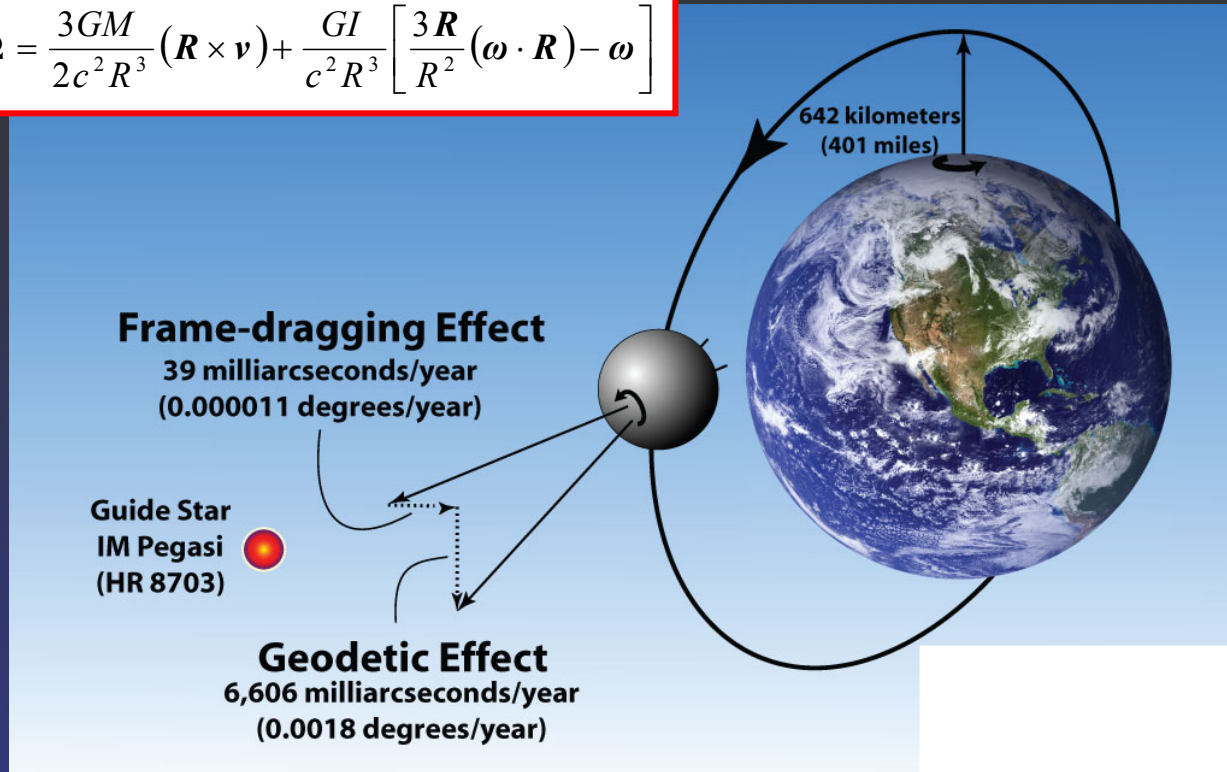
HEPL-AA Seminar

June 17, 2009



The Relativity Mission Concept

$$\Omega = \frac{3GM}{2c^2 R^3} (\mathbf{R} \times \mathbf{v}) + \frac{GI}{c^2 R^3} \left[\frac{3\mathbf{R}}{R^2} (\boldsymbol{\omega} \cdot \mathbf{R}) - \boldsymbol{\omega} \right]$$



- **Geodetic Effect**
 - ◆ Space-time curvature ("the missing inch")
- **Frame-dragging Effect**
 - ◆ Rotating matter drags space-time ("space-time as a viscous fluid")

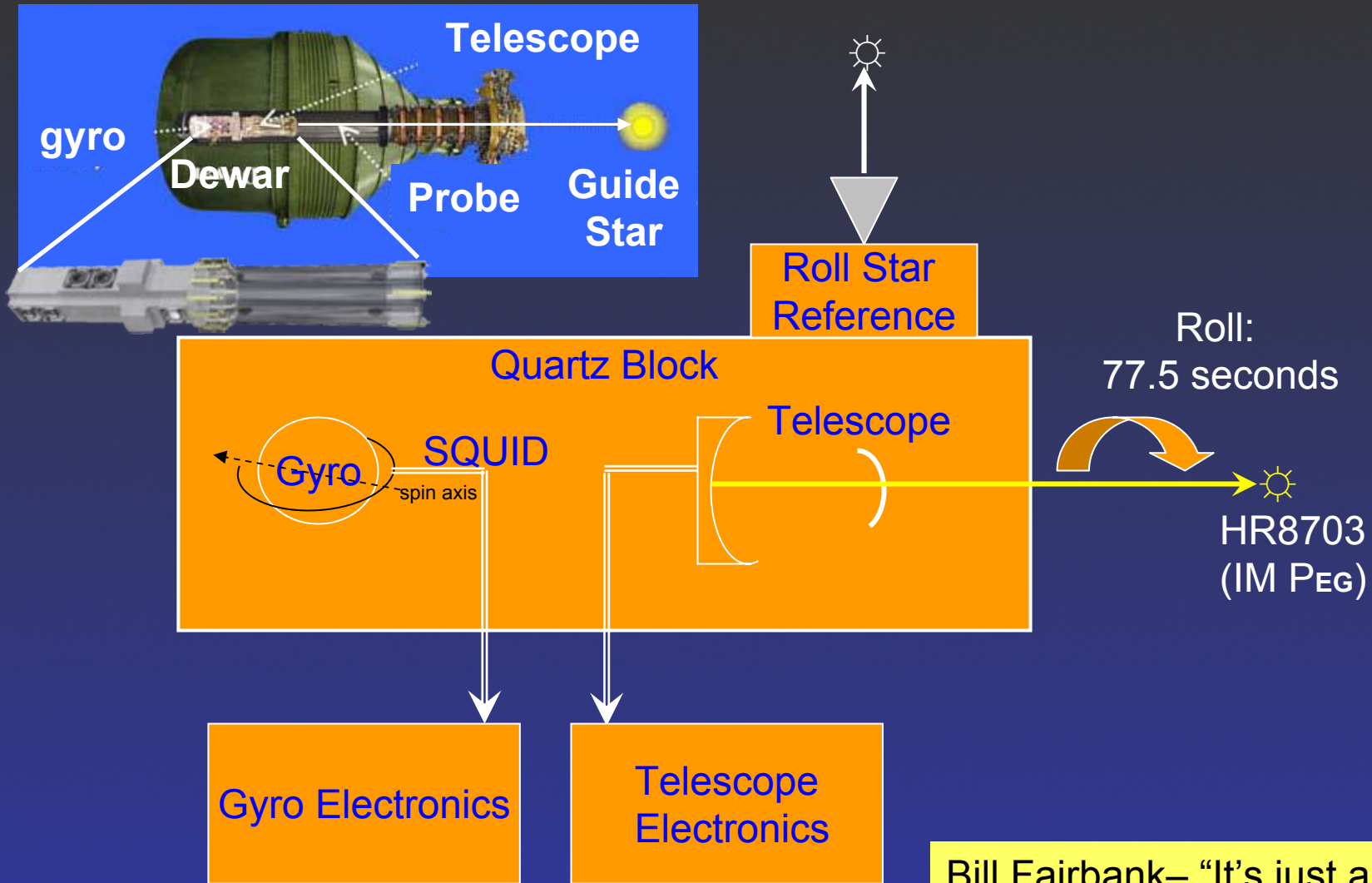
Launch: April 20, 2004 – 09:57:24



Stanford Mission Operations Center



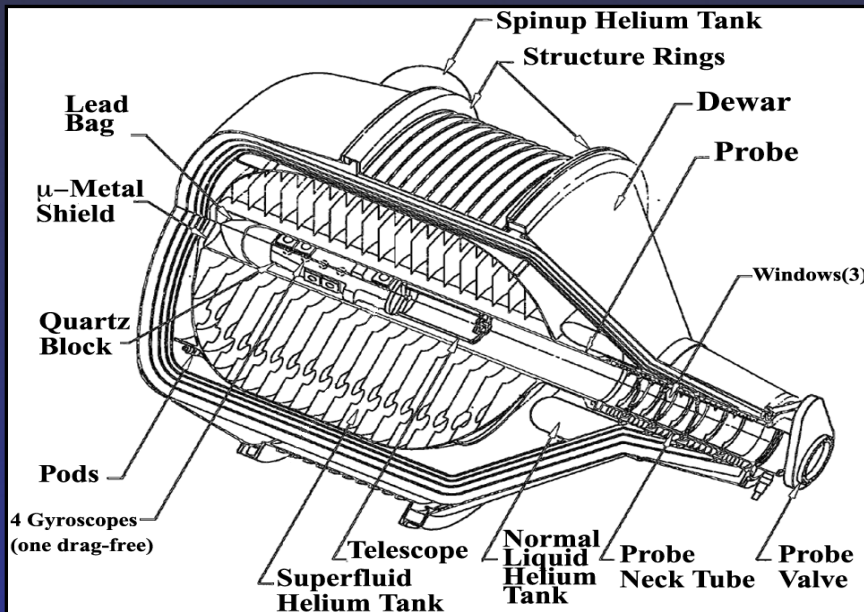
Measurement of Gyroscope Orientation Relative to Position of Guide Star



Bill Fairbank— “It’s just a star, a telescope, & a spinning sphere”

The GP-B Challenge

- ♦ Gyroscope (G) 10^7 times better than best 'modeled' inertial navigation gyros
- ♦ Telescope (T) 10^3 times better than best prior star trackers
- ♦ Gyro Readout \rightarrow calibrated to parts in 10^5
- ♦ G – T \rightarrow <1 marc-s subtraction within pointing range



Basis for 10^7 advance
in gyro performance

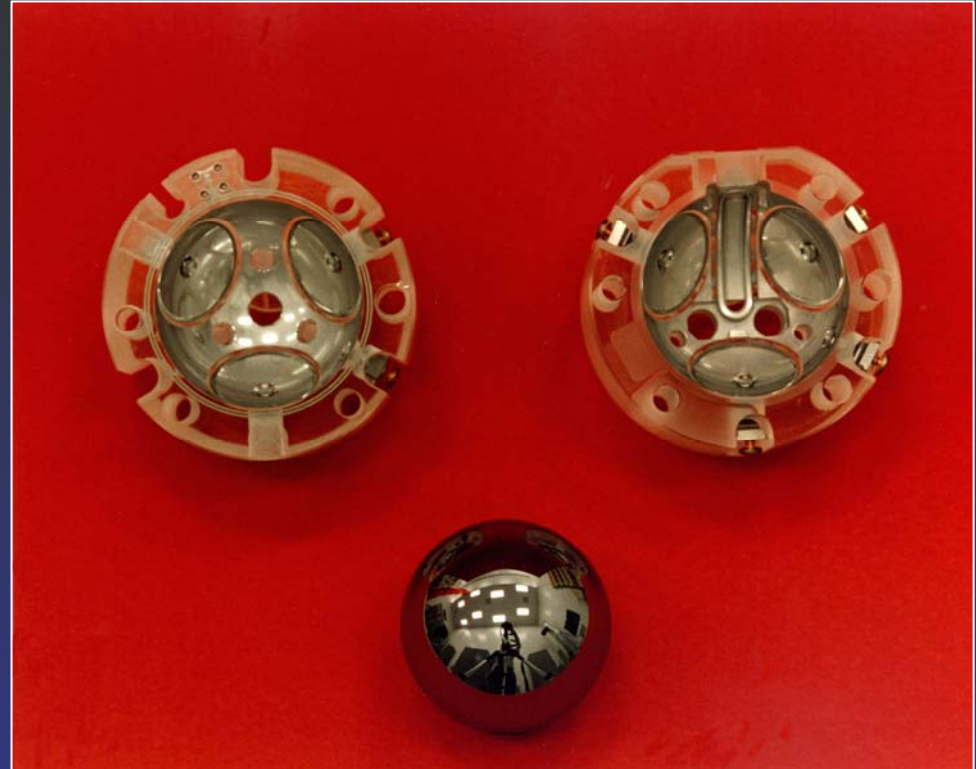
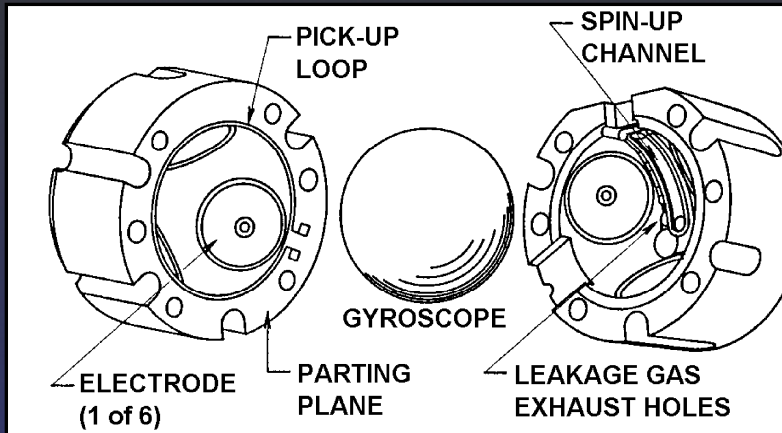
Space

- reduced support force, "drag-free"
- roll about line of sight to star

Cryogenics

- magnetic readout & shielding
- thermal & mechanical stability
- ultra-high vacuum technology

The GP-B Gyroscope



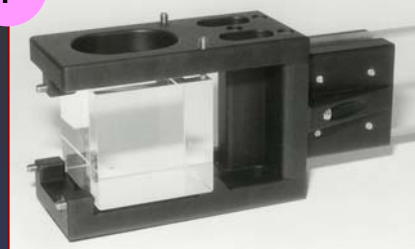
- **Electrical Suspension**
- **Gas Spin-up**
- **Magnetic Readout**
- **Cryogenic Operation**

Challenge 1: $< 10^{-11}$ deg/hr Classical Drift

Seven Near Zeros

- "Drag-free" (cross track) $< 10^{-11}$ g met
- Rotor inhomogeneities $< 10^{-6}$ met
- Rotor asphericity < 10 nm met
- Magnetic field $< 10^{-6}$ gauss met
- Pressure $< 10^{-12}$ torr met
- Electric charge $< 10^8$ electrons met
- **Electric dipole moment 0.1 V-m issue**

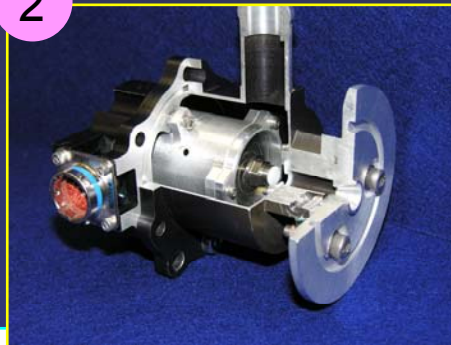
1



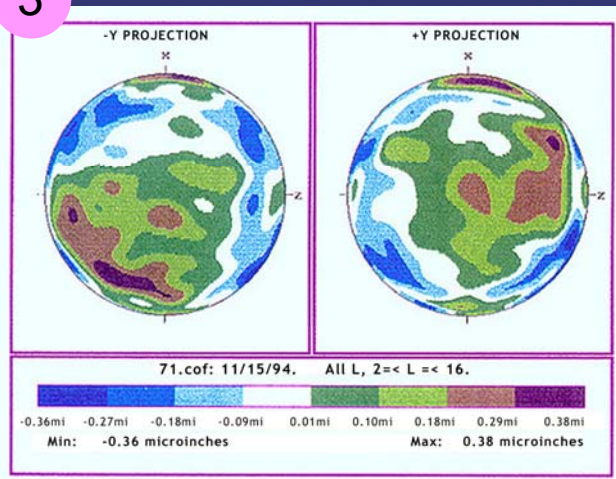
4



2



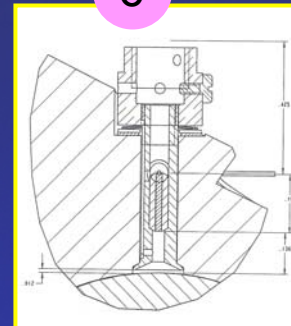
3



5

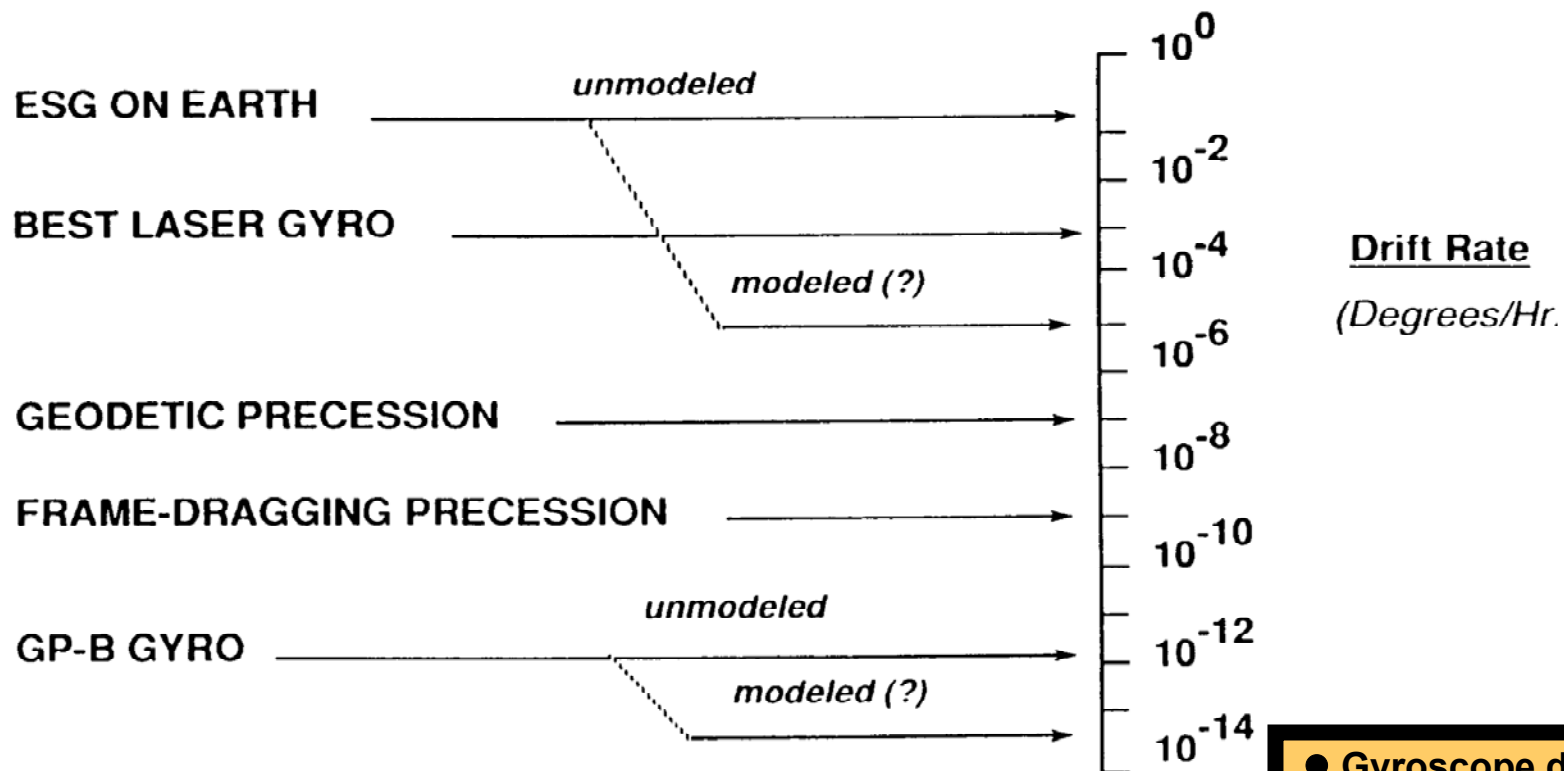


6



GP-B Performance

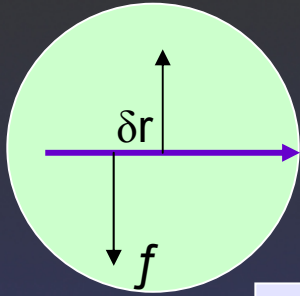
Why go to space?



Space improves gyro accuracy
by $> 50,000,000!$

- Gyroscope drift ≤ 0.05 marcsec/yr
- Readout error effect ≤ 0.08 marcsec/yr
- Guide star uncertainty ≤ 0.09 marcsec/yr

Mass-Unbalance & Drag-Free



Drift-rate

$$\Omega = T / I\omega_s$$

Torque

$$T = M f \delta r$$

Moment of Inertia

$$I = 2/5 M r^2$$

requirement $\Omega < \Omega_0 \sim 0.1 \text{ marc-s/yr}$ ($1.54 \times 10^{-17} \text{ rad/s}$)

$$f \frac{\delta r}{r} < \frac{2}{5} v_s \Omega_0$$

$$v_s = \omega_s r = 950 \text{ cm/s} \quad (80 \text{ Hz})$$

On Earth ($f = g$)

$$\frac{\delta r}{r} < 5.8 \times 10^{-18} \quad (\text{ridiculous})$$

Standard satellite ($f \sim 10^{-8} \text{ g}$)

$$\frac{\delta r}{r} < 5.8 \times 10^{-10} \quad (\text{unlikely})$$

GP-B drag-free ($f \sim 10^{-12} \text{ g}$
cross-track average)

$$\frac{\delta r}{r} < 5.8 \times 10^{-6} \quad (\text{straightforward})$$

All GP-B rotors mass unbalance
< 10 nm (2 cm radius) or

$$\frac{\delta r}{r} \sim \underline{5 \times 10^{-7}}$$

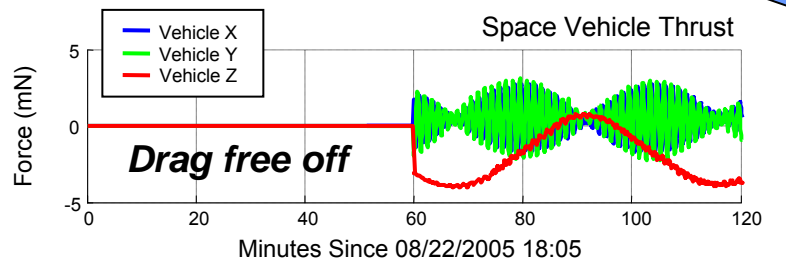
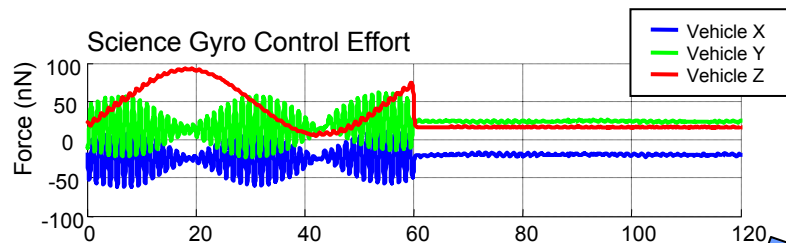
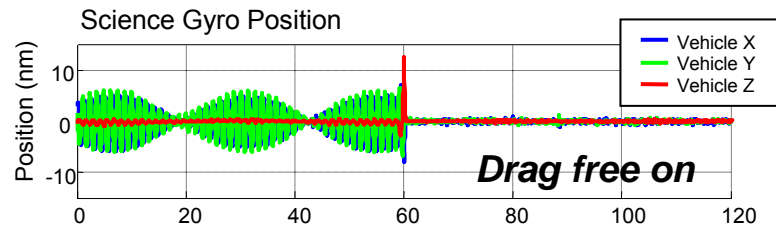
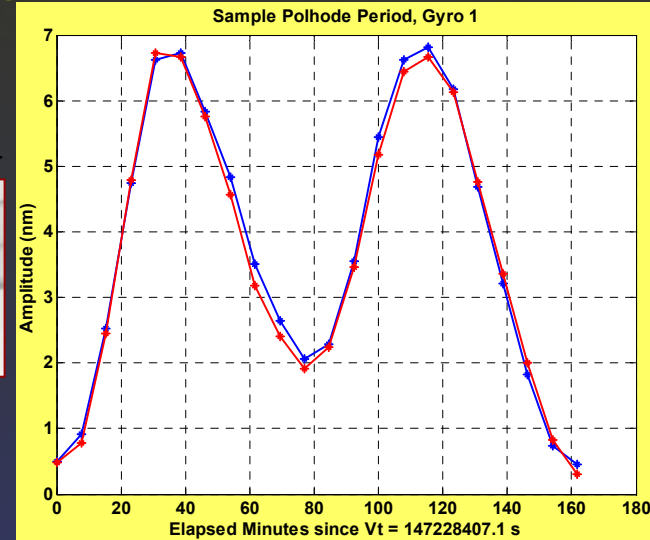
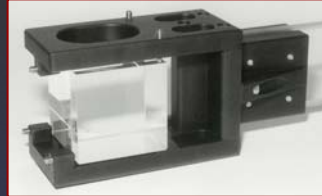
drift-rate for the
drag-free GP-B
< 0.01 marc-s/yr

Drag-free eliminates mass-unbalance torque -- and key to understanding/quantification of other support torques

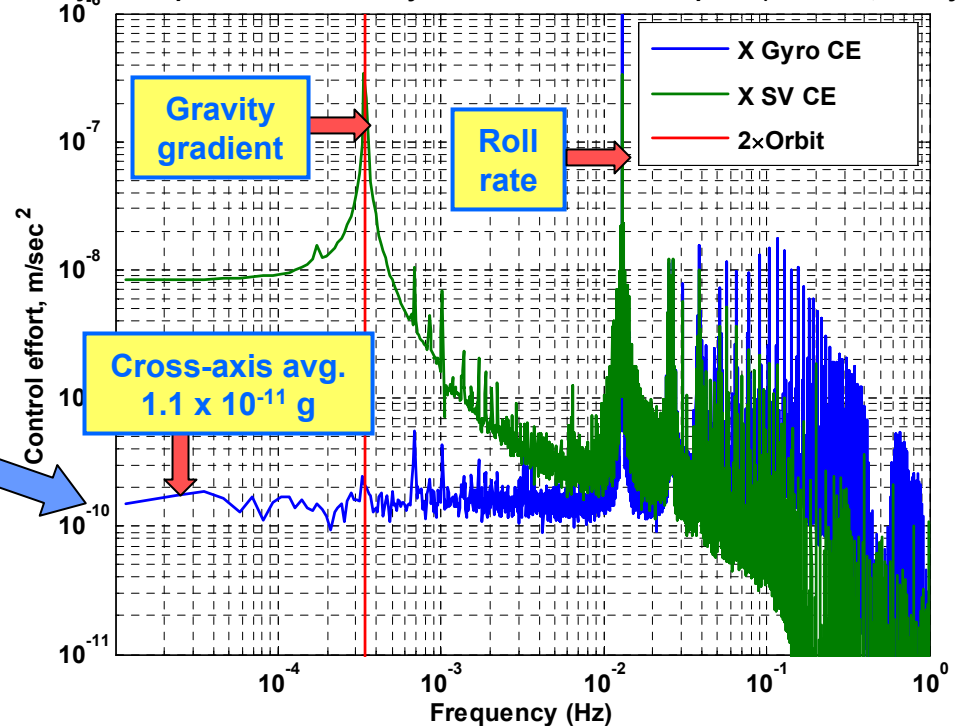
Mass Unbalance & Drag Free On-orbit Results

Rotor Mass Unbalance < 10 nm for all gyros

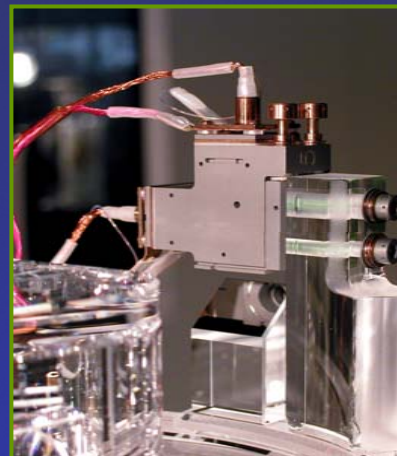
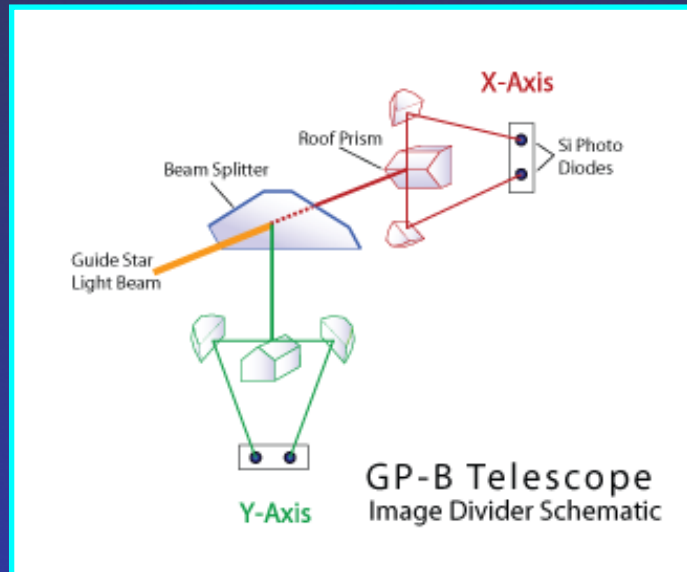
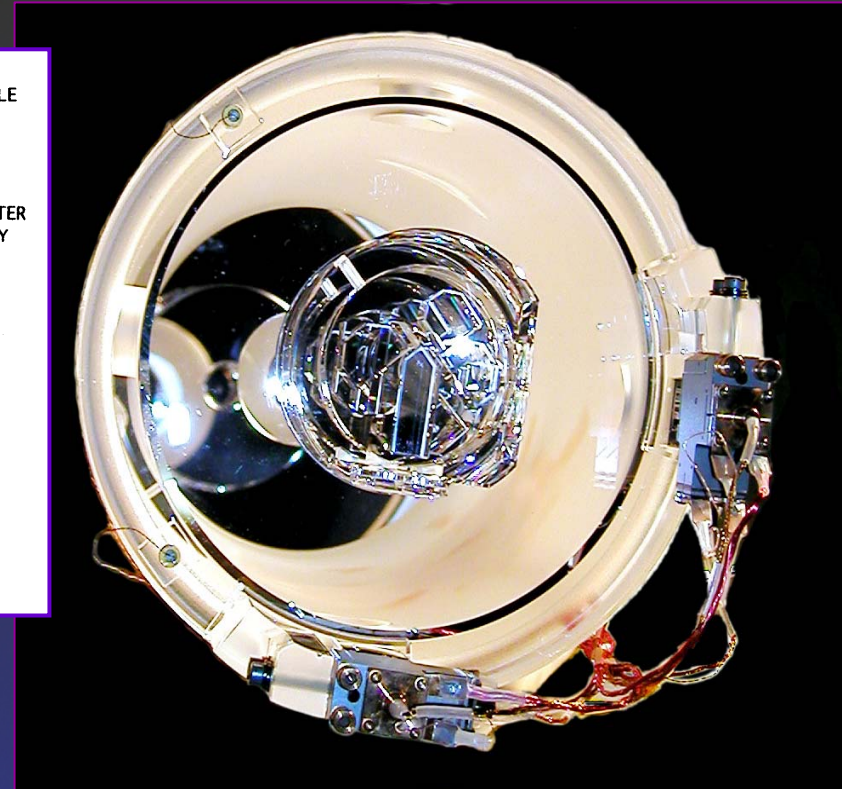
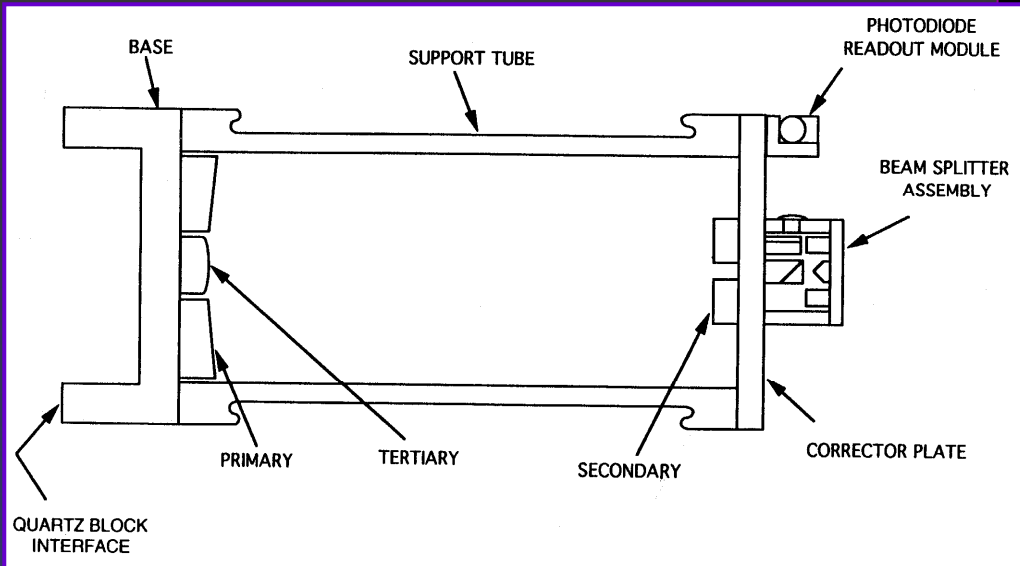
Drag Free (cross-track 10^{-11} g)



Gyro 1 - Space Vehicle and Gyro Ctrl Effort - Inertial space (2005/200, 14 days)



Challenge 2: Sub-milliarc-s Star Tracker



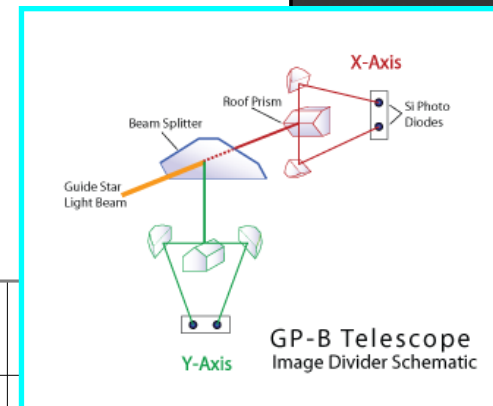
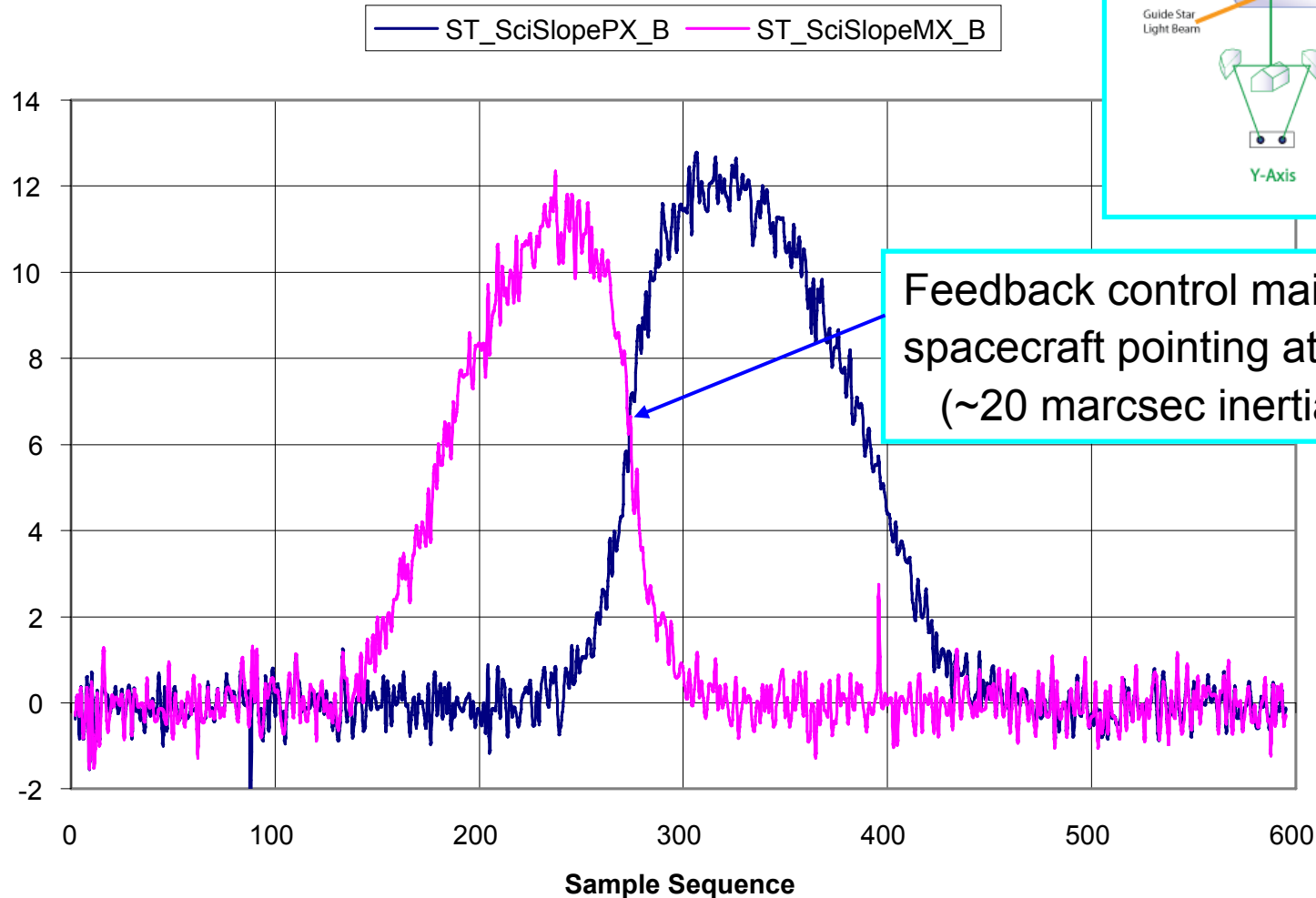
Detector Package



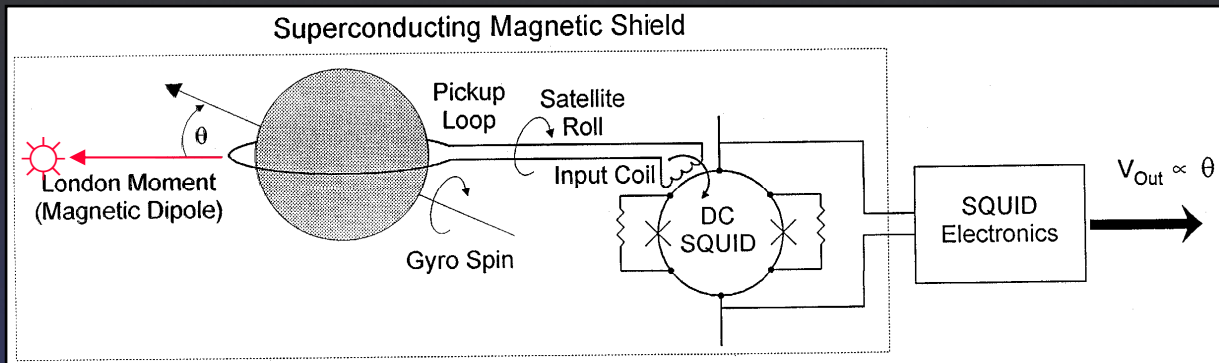
Dual Si Diode Detector

Star Tracking Telescope: On-Orbit

Telescope Detector Signals
from IM Peg Divided by Rooftop Prism



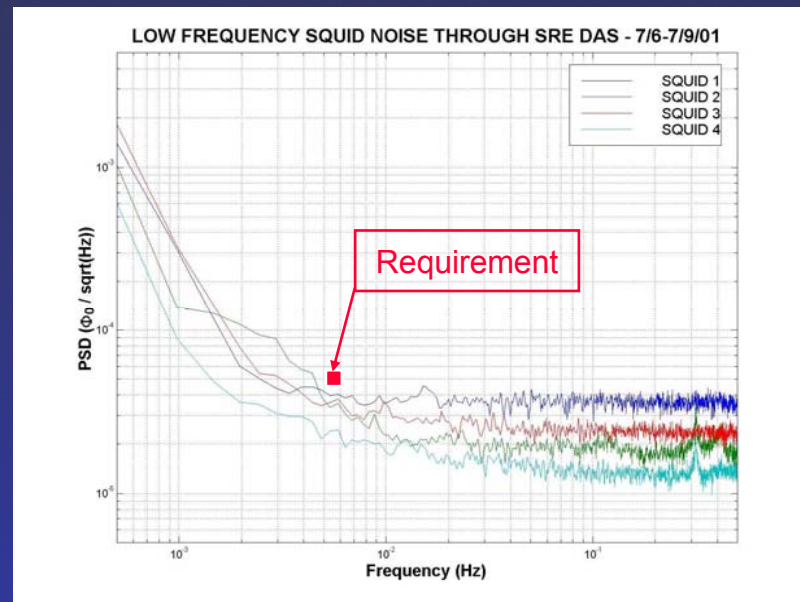
Challenge 3: Gyro Readout



"SQUID" → 1 marc-s in 5 hours

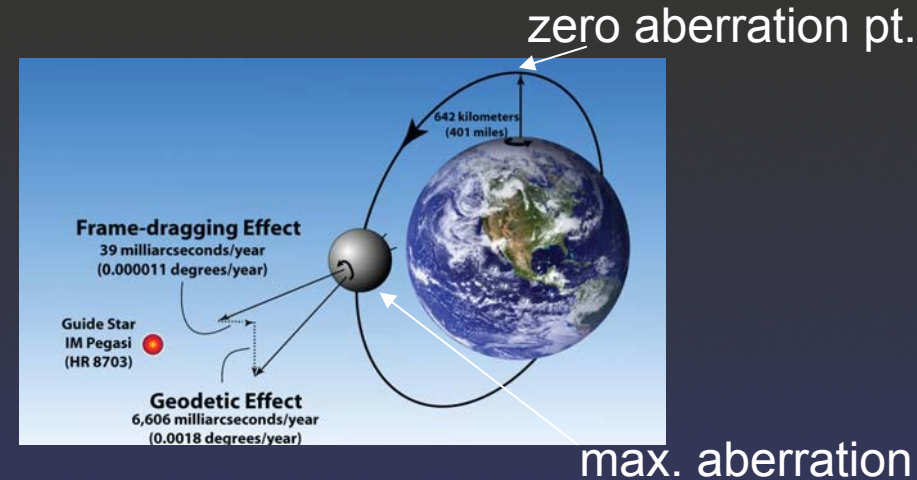
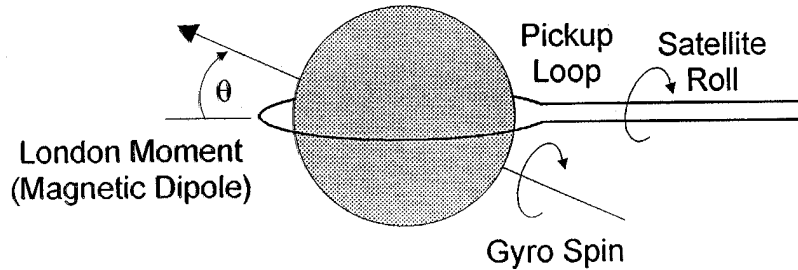
4 Requirements/Goals

- ◆ SQUID noise 190 marc-s/ $\sqrt{\text{Hz}}$
- ◆ Centering stability < 50 nm
- ◆ DC trapped flux < 10^{-6} gauss
- ◆ AC shielding > $\sim 10^{12}$



Challenge 3: Gyro Readout Calibration

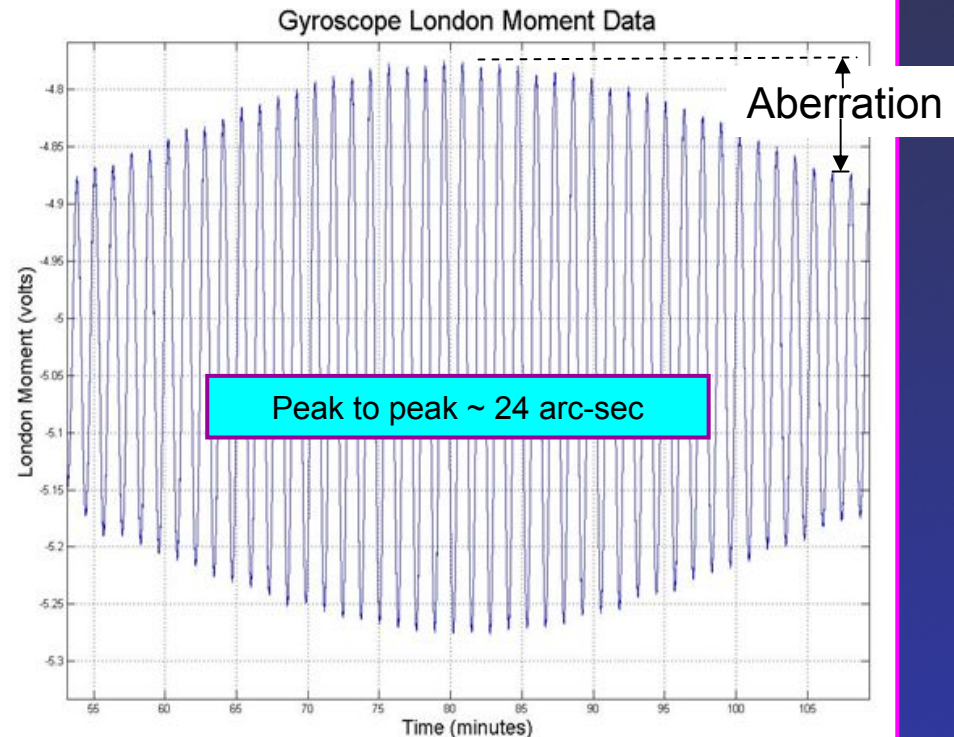
Measurement System



Aberration: A Natural Calibration

Orbital motion \Rightarrow varying apparent position of star
 Cause: transverse velocity of telescope to starlight
 (v_{orbit}/c + special relativity correction)

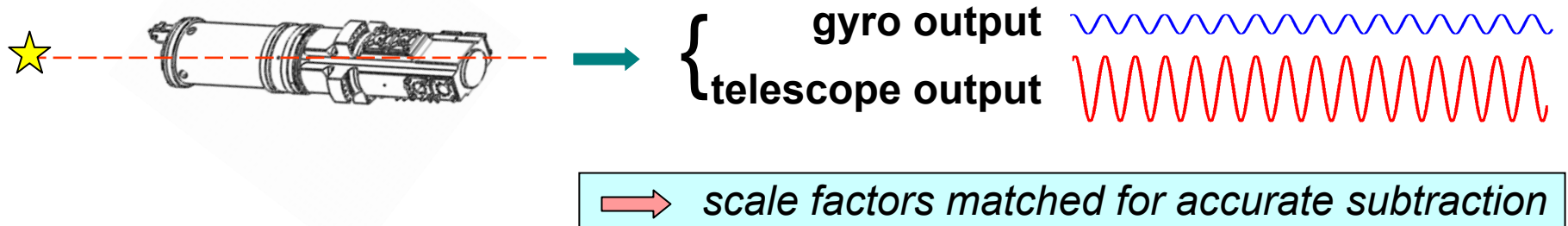
S/V around Earth -- 5.1856 arc-s @ 97.5-min period
 Earth around Sun -- 20.4958 arc-s @ 1-year period



Challenge 4: Gyro-Telescope Matching

How do we subtract imperfect telescope pointing from the gyro signal?

Dither -- Slow 60 marc-s oscillations injected into pointing system



Scale factor matching allows G-T subtraction to < 1 marcsec

Actualities of GP-B

Good

- **gyroscopes**
 - ♦ 10^5 times better than best inertial navigation gyroscopes
 - ♦ SQUID noise < expected
 - ♦ Trapped magnetic flux meets spec
 - ♦ excellent charge control & centering stability
 - ♦ τ 's ~ 7200 to 26,100 years
- **telescope**
 - ♦ superb overall performance
- **dewar**
 - ♦ beats design hold time
- **orbit within 100 m of ideal**

Less than ideal

- *polhode rate variation*
 - *misalignment torque*
 - *resonance torque*
-
- *segmented data*
 - *interference from ECU*
 - *out-of-spec pointing*

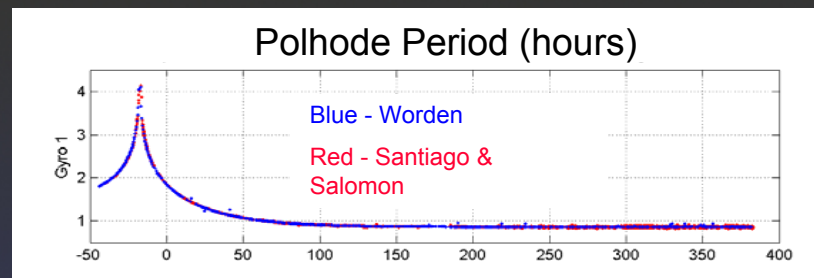
New Challenge

Systematics & data grading

Challenge 5: Data Analysis Subtleties

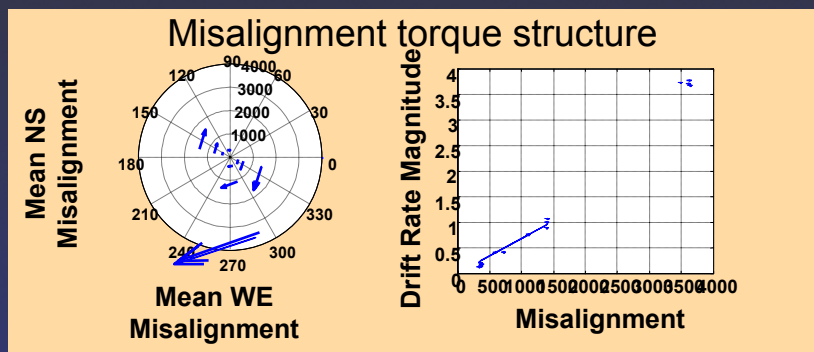
A. Polhode rate variations affect scale factor (C_g) determinations

- Discovered in early science phase
- Alex Silbergleit talk July 1



B. Misalignment torques

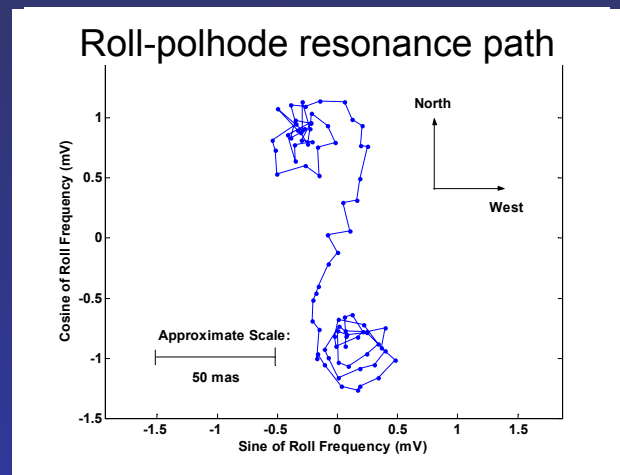
- Discovered in post-science calibration phase
- Mac Keiser talk next week



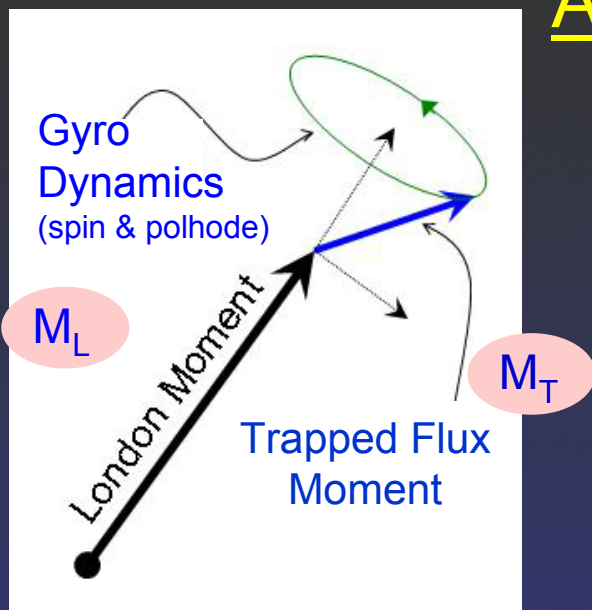
C. Roll-polhode resonance torques

- Discovered during data reduction phase
- Mac Keiser talk next week

All due to one physical cause (patch effect)



Evolving Polhode Impact on Actual London Moment Readout



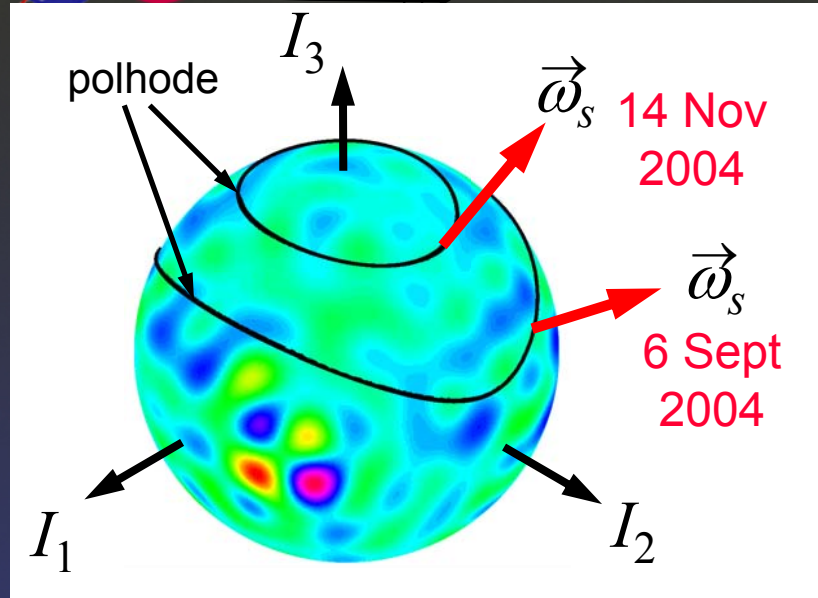
London field at 80 Hz: 57.2 μG

Trapped fields	Gyro 1	3.0 μG
	Gyro 2	1.3 μG
	Gyro 3	0.8 μG
	Gyro 4	0.2 μG

Trapped flux contributes to readout scale factor

- Expected prior to launch
- Adds to London Moment scale factor
- Trapped flux scale factor modulated by body dynamics (polhode)
- Evolving polhode prevents averaging or other simple analysis technique

Trapped Flux Mapping: C_g Determination



TFM determines evolving polhode phase to 1° over the full mission

- Fully resolves gyro scale factor
- Crucial input for torque analysis



Alex Silbergleit

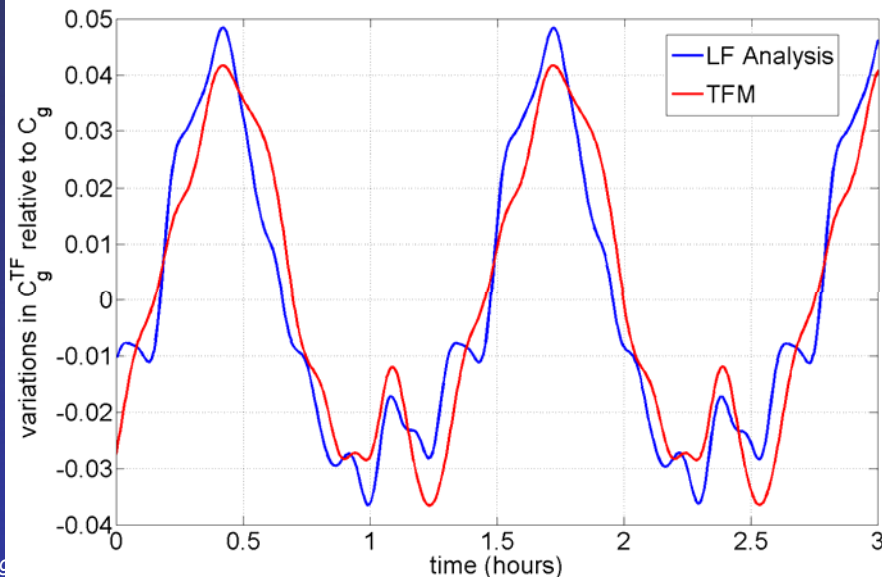


John Conklin
(Ballhaus AA award)

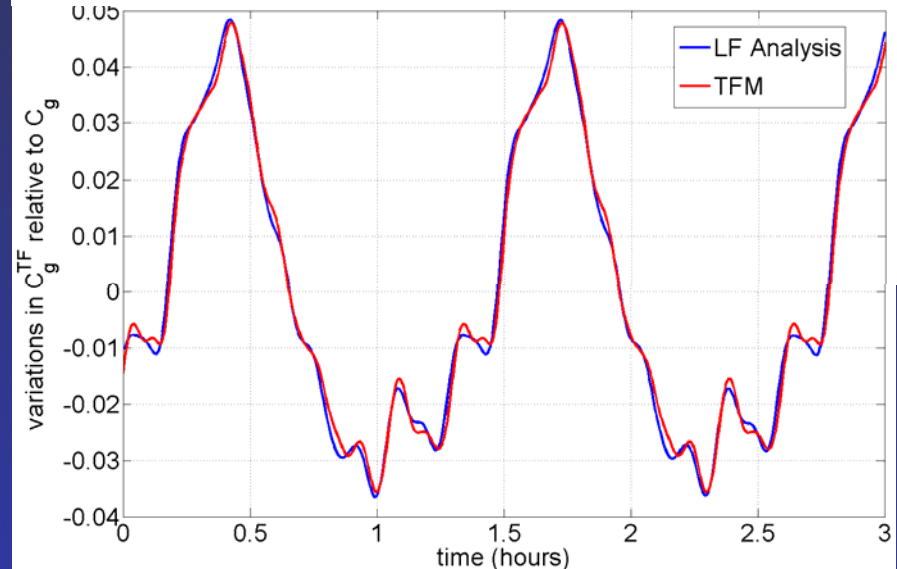


Mac Keiser

Nov. 2007, Gyro 1, Fit residuals = 14%



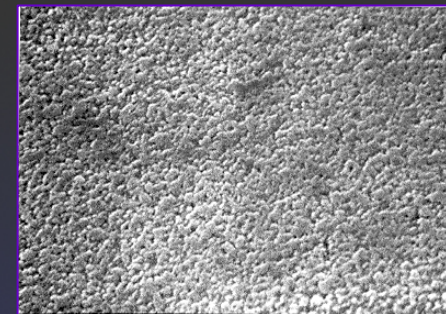
Aug. 2008, Gyro 1, Fit residuals = 1%



The GPB 'Patch Effect' History

Pre-launch investigation

- ◆ Rotor electric dipole moment + field gradient in housing
- ◆ 100 mV contact potentials mitigated by minute grain size, $0.1 \mu\text{m} \ll 30 \mu\text{m}$ rotor-electrode gap
- ◆ Kelvin probe measurements on flat samples



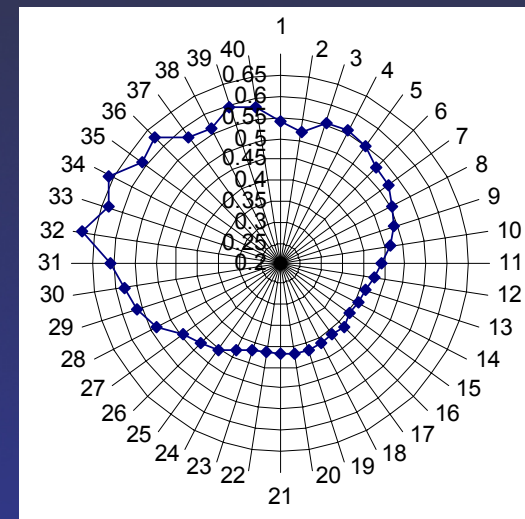
SEM image of rotor Nb film
average grain size $0.1 \mu\text{m}$

On-orbit discoveries (Sasha Buchman talk next week)

- ◆ *Polhode damping* (July 2004)
- ◆ *Drag-free z acceleration* (Sept. 2004)
- ◆ *Spin down rate > gas damping* (Feb. 2005)
- ◆ *Misalignment torques* (Aug. 2005)
- ◆ *Roll-polhode resonance torques* (Jan. 2007)

Post-launch ground-based investigations

- ◆ Work function profile via UV photoemission
- ◆ Detailed analytical modeling



Work function polar plot



Origin of Misalignment & Resonance Torque

1. Expand rotor & housing potentials

- Using spherical harmonics

2. Derive stored energy

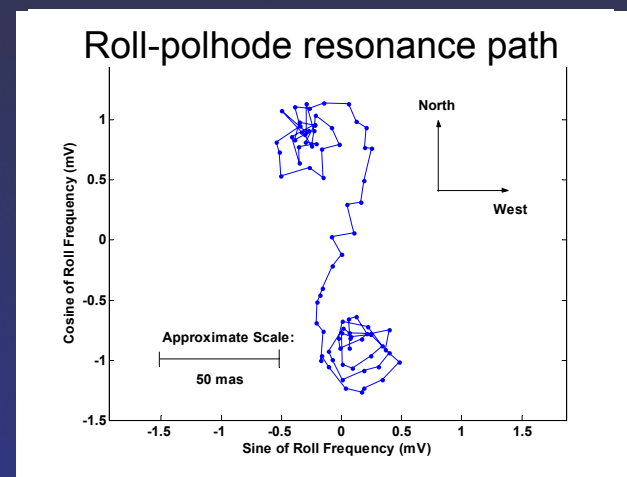
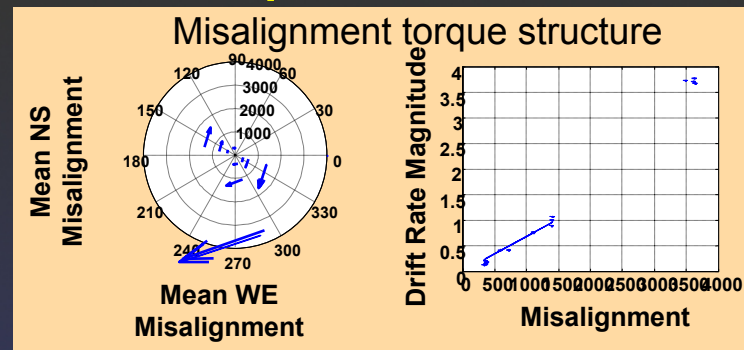
- Laplace's equation
- Between rotor & housing

3. Find torque: Derivative of the energy

- WRT angles defining the spin orientation
- Spin average

- Roll ave. torque proportional & perpendicular to misalignment
- Non roll averaged torque when polhode harmonic = roll freq.

These torques follow directly from randomly distributed patch potentials



Treating Misalignment Torque

The 2 Methods

Algebraic: Filtering machinery to explicitly model torques

→ provides separation from relativity

Geometric: Plot rates against misalignment phase ϕ

→ component of relativity free of misalignment torques

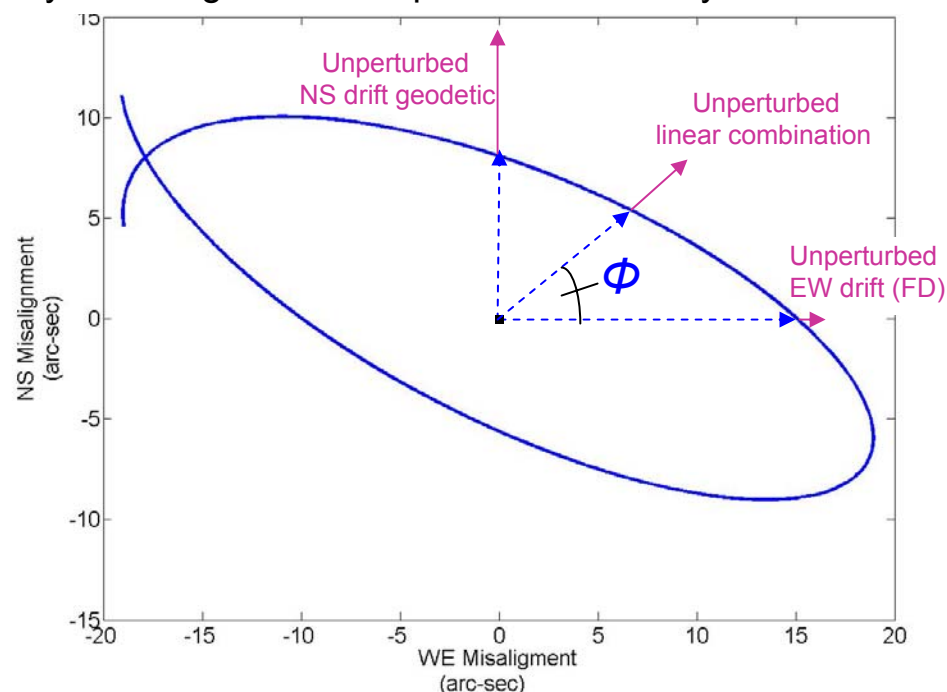
A Geometrical Insight

- Drift rate free of torque parallel to misalignment
- Annual aberration alters torque direction
 - Over time provides geodetic & FD measurement
- A truly *physical* modeling process



Mac Keiser

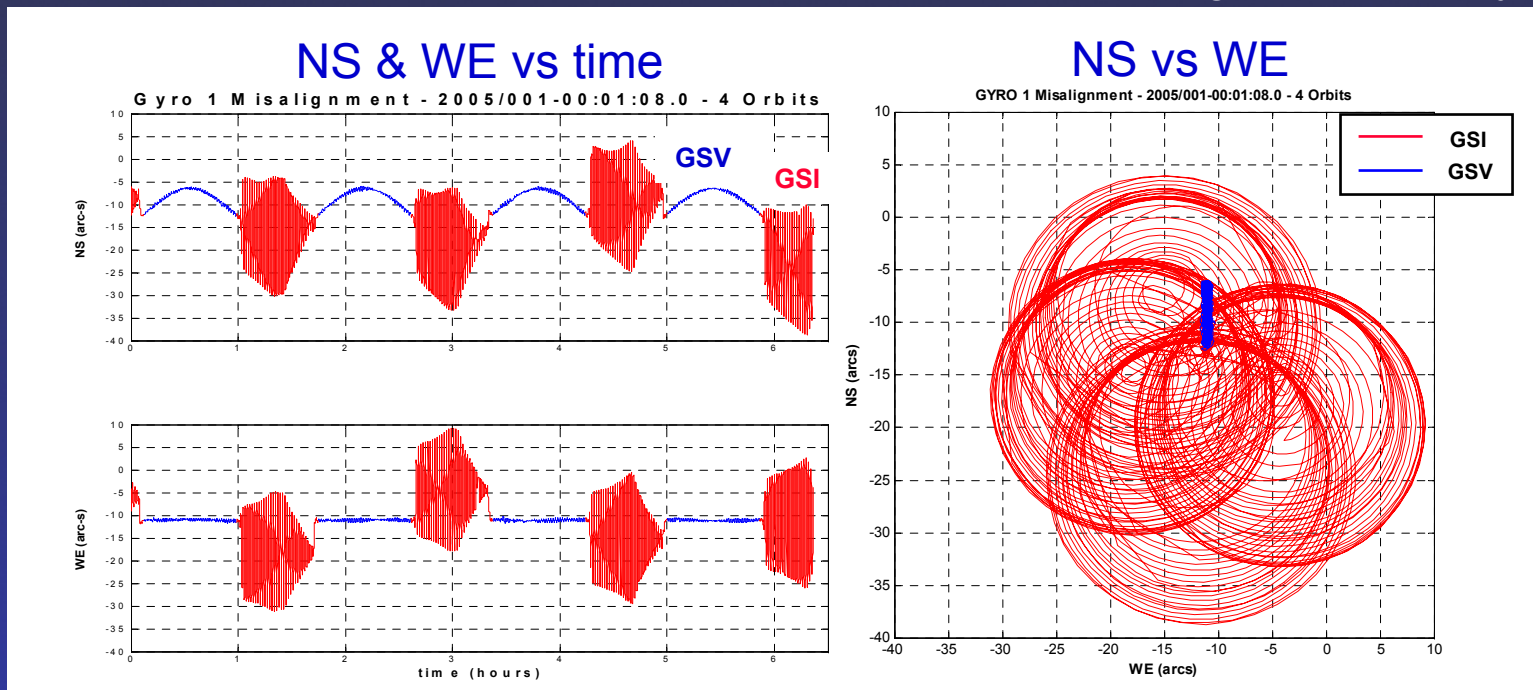
Gyro misalignment & unperturbed relativity measurement



Obtaining Continuous Misalignment History

- Required to remove effect of misalignment torque
- Science gyroscopes provide precision misalignment information when guide star occulted

Continuous Guide-Star Valid / Guide-Star Invalid misalignment history



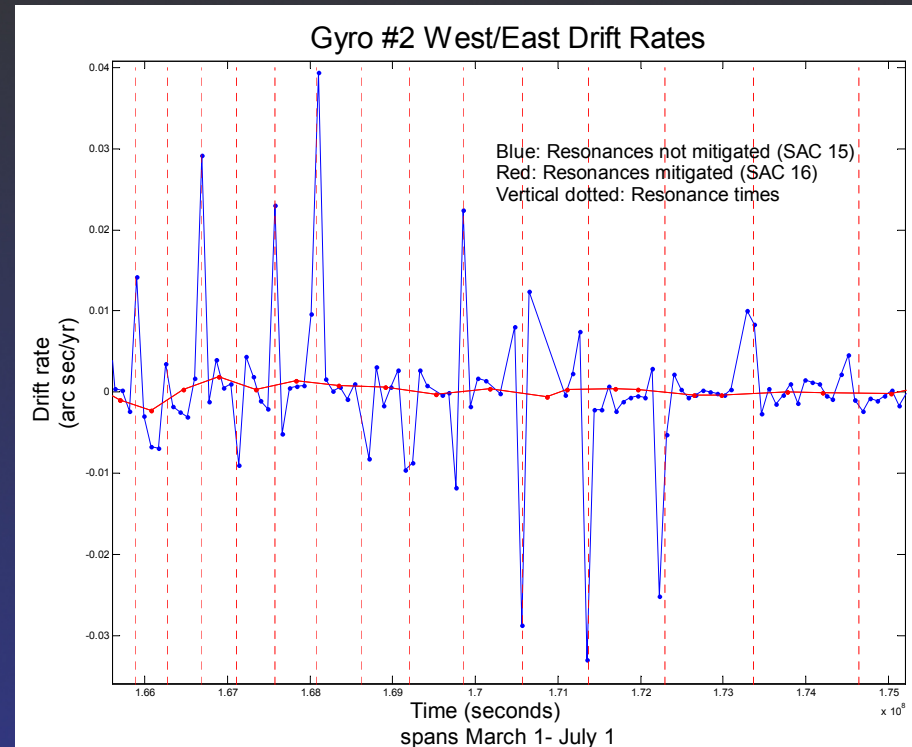
Early Methods for treating Resonance Torque

1. Excise data during resonances

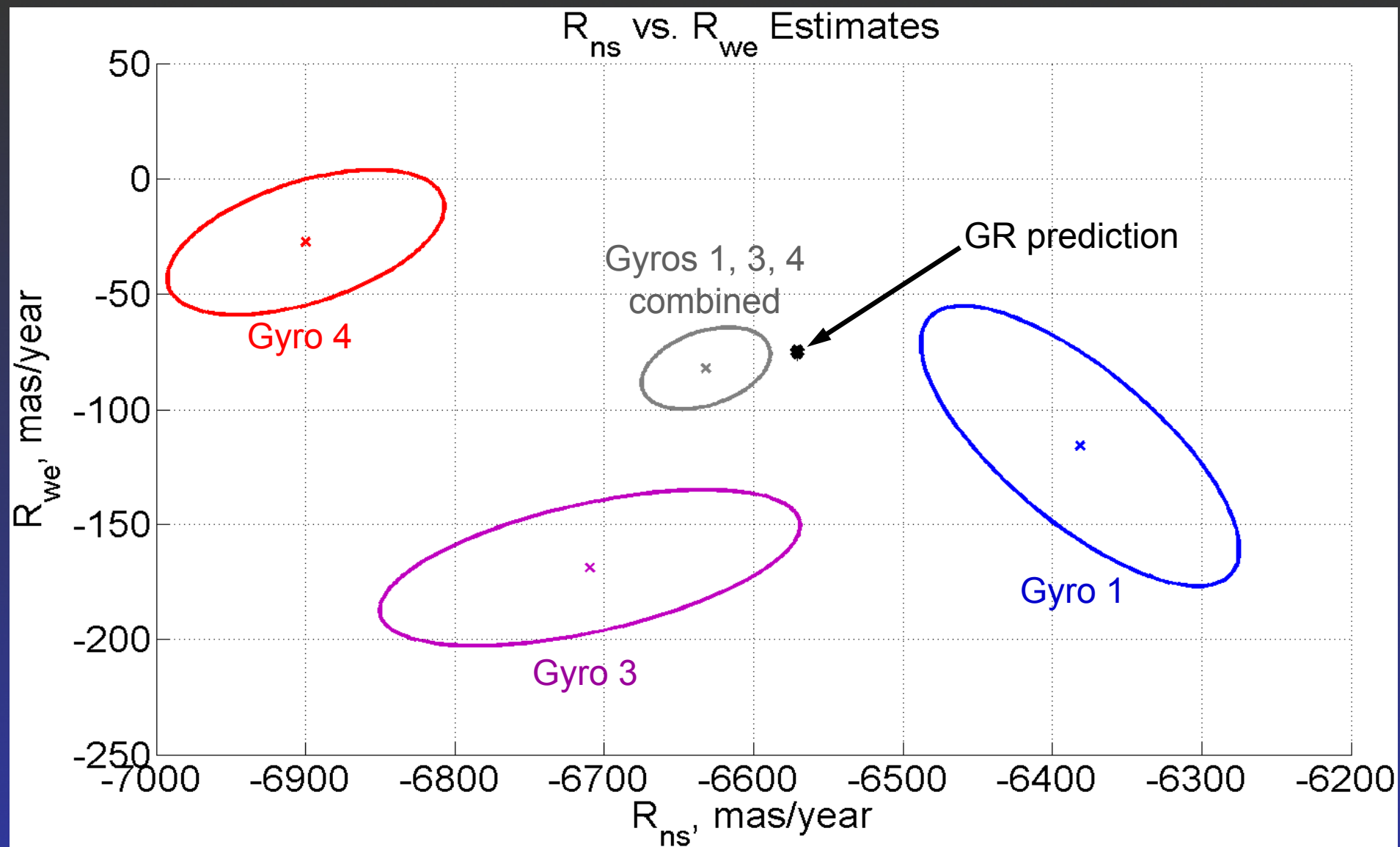
- Blue curve shows large drift during resonances
- Red curve eliminates spikes by excising data

2. Model drift during resonances

- A unique drift rate during each
- A separate drift rate elsewhere

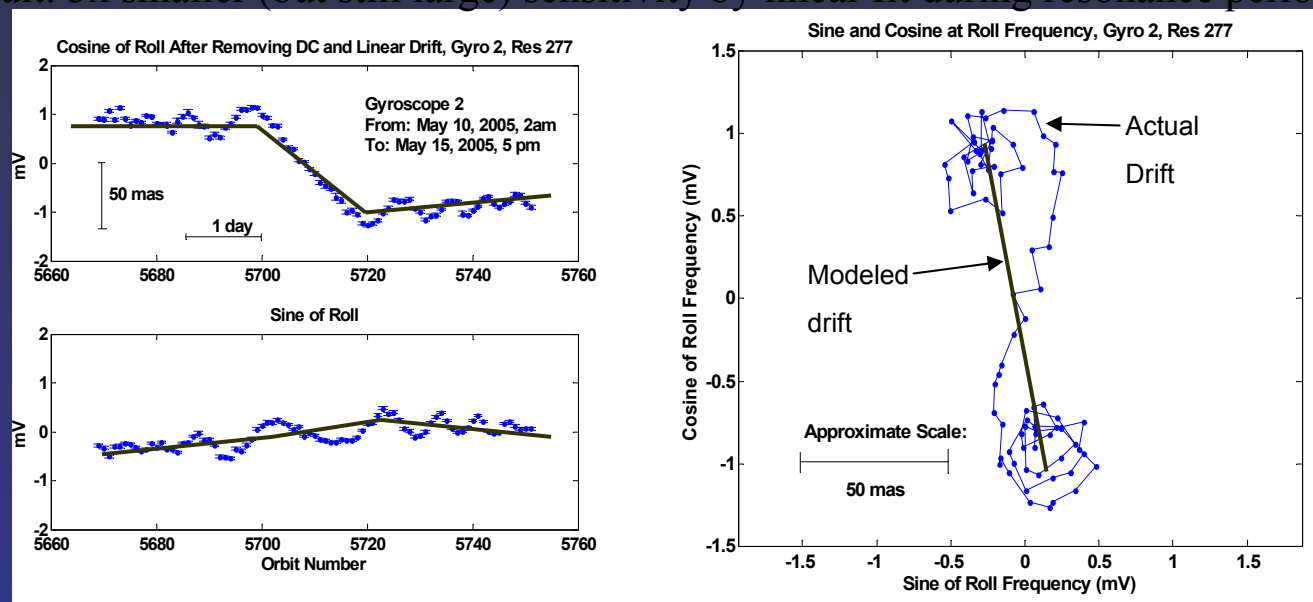


Early Results [Nov '07]



Resonance Torque Modeling Limitations


- Issue: High sensitivity for resonance torque modeling (2007)
 - Leading term in experiment error (relativity rate scatter)
 - First result based upon excluding data within resonance periods
 - Second result: 3x smaller (but still large) sensitivity by linear fit during resonance periods

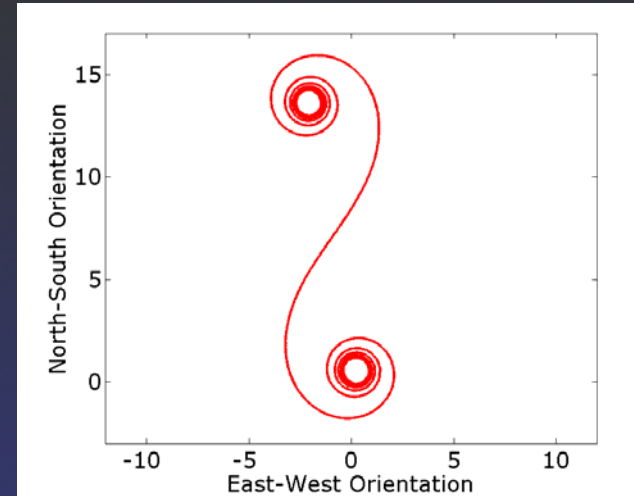


Next step: Once per orbit model improves sensitivity & experiment error

Improved Resonance Torque Modeling

- Path predicted from rotor & housing potentials

- Roll averaging fails when $\omega_r = n\omega_p$
- Orientations follow Cornu spiral 
- Magnitude & direction depend on patch distribution, roll & polhode phases at resonance



- Example: Gyro 2, Resonance 277 – Oct 25, '04



Jeff Kolodziejczak



Mac Keiser



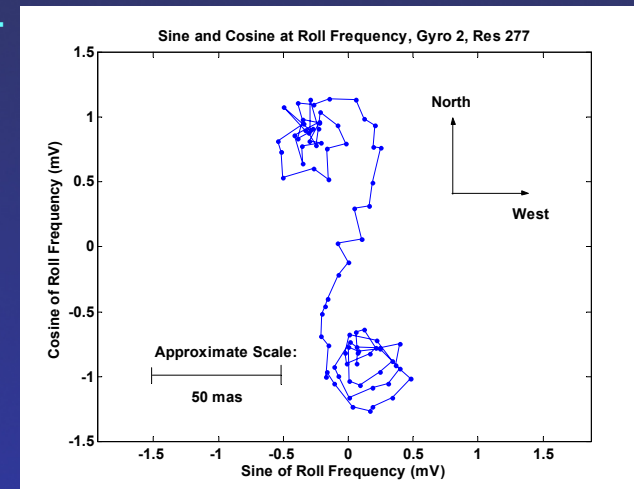
Alex Silbergleit



Michael Heifetz



Vladimir Solomonik



The Science Equations [Aug 2008]

- Add resonance torque model to equations of motion
 - Follows directly from randomly distributed patch potentials

$$\begin{aligned}\frac{ds_{NS}}{dt} &= r_{NS} + k(t)\mu_{EW} + \sum A_m(t) \cos \Delta\phi_m - B_m(t) \sin \Delta\phi_m \\ \frac{ds_{EW}}{dt} &= r_{EW} - k(t)\mu_{NS} + \sum A_m(t) \sin \Delta\phi_m + B_m(t) \cos \Delta\phi_m\end{aligned}$$

Relativity

Misalignment torque

Resonance torque

- Resonance condition

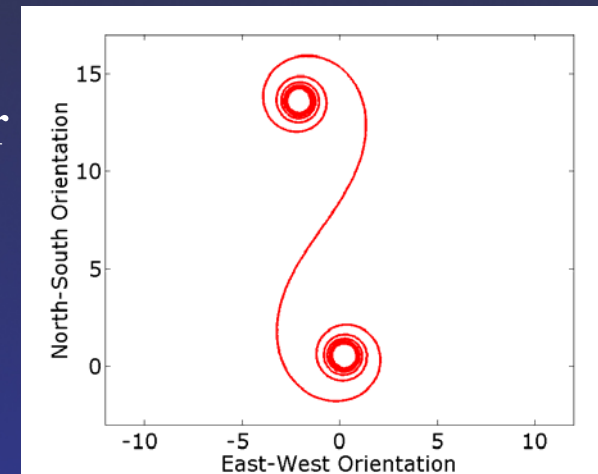
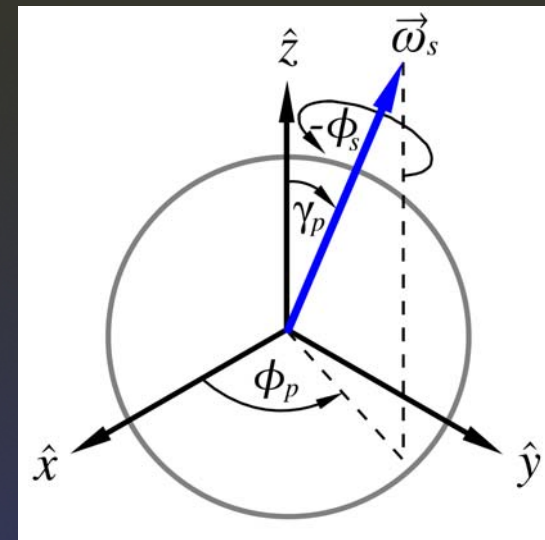
$$\Delta\phi_m = \phi_{\text{roll}} - m\phi_p$$



Treatment of roll-polhode term hinges on TFM

Resonance Torque Modeling Implementation

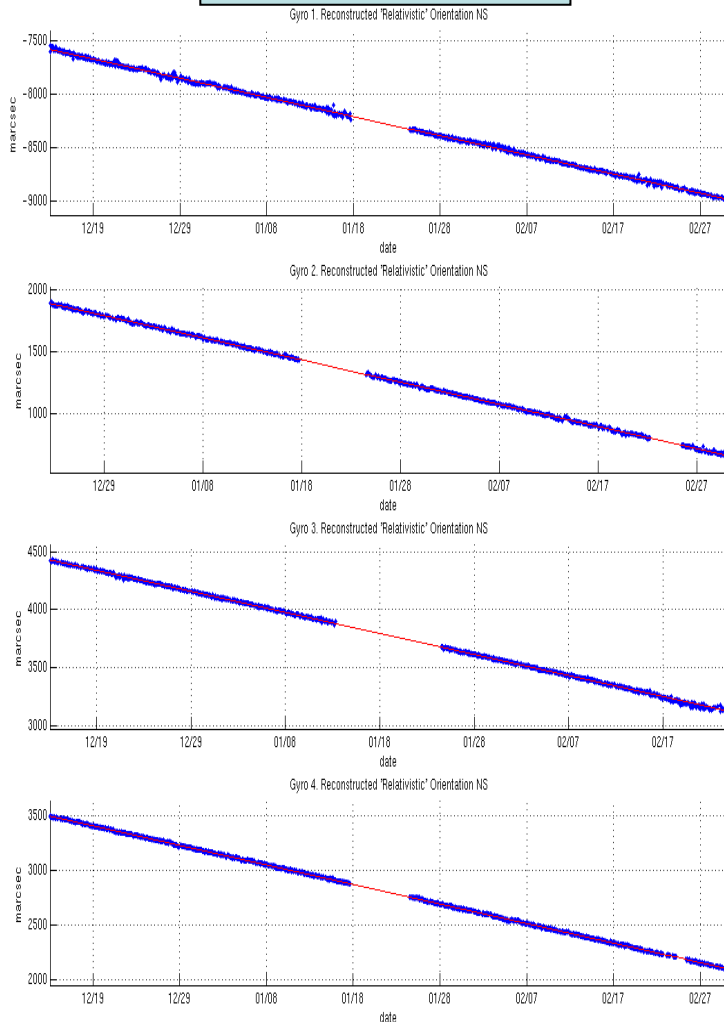
- Torque coefficients tied to gyro body
- TFM provides accurate body dynamics profile
 - $\gamma_p \phi_p$ (in addition to readout scale factor)
 - TFM: body fixed trapped flux as marker
- Torque model uses TFM polhode
- Requires accuracy for tracking high order polhode harmonics



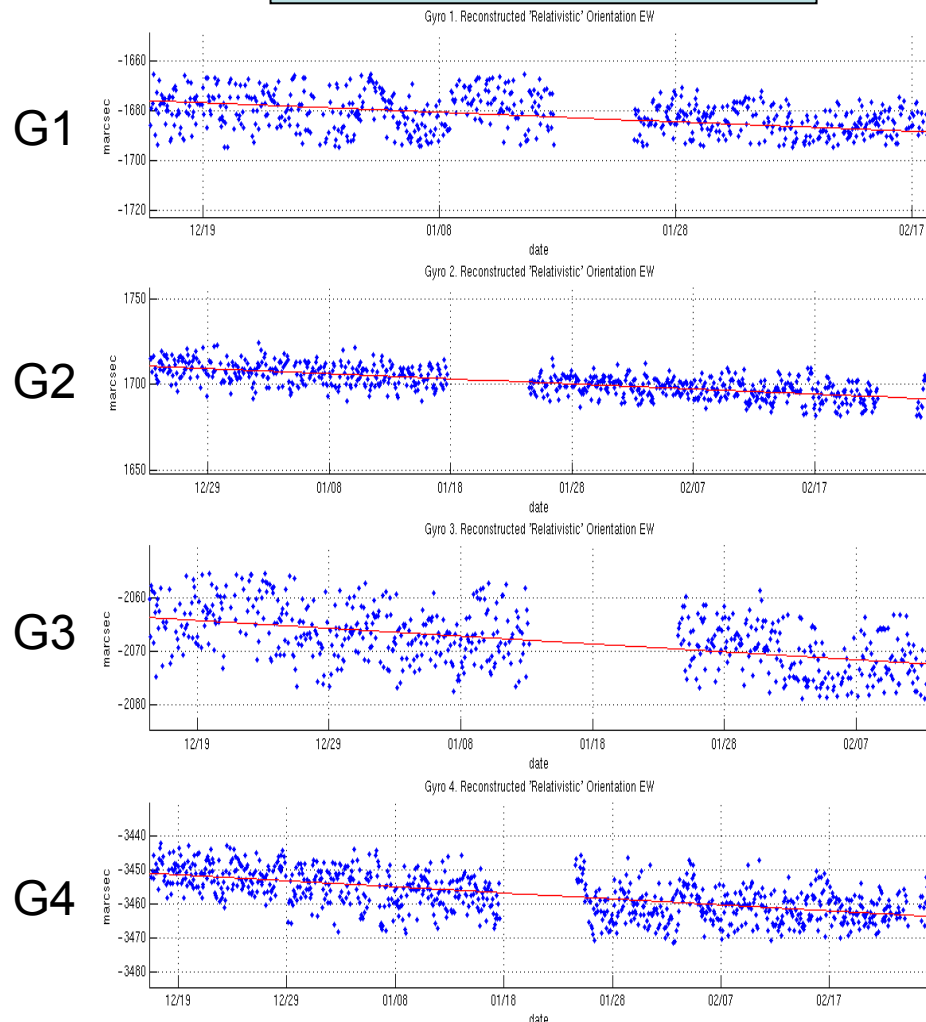
Body dynamics determined from TFM is KEY

Full Model Results as of Dec '08

N-S (Geodetic)

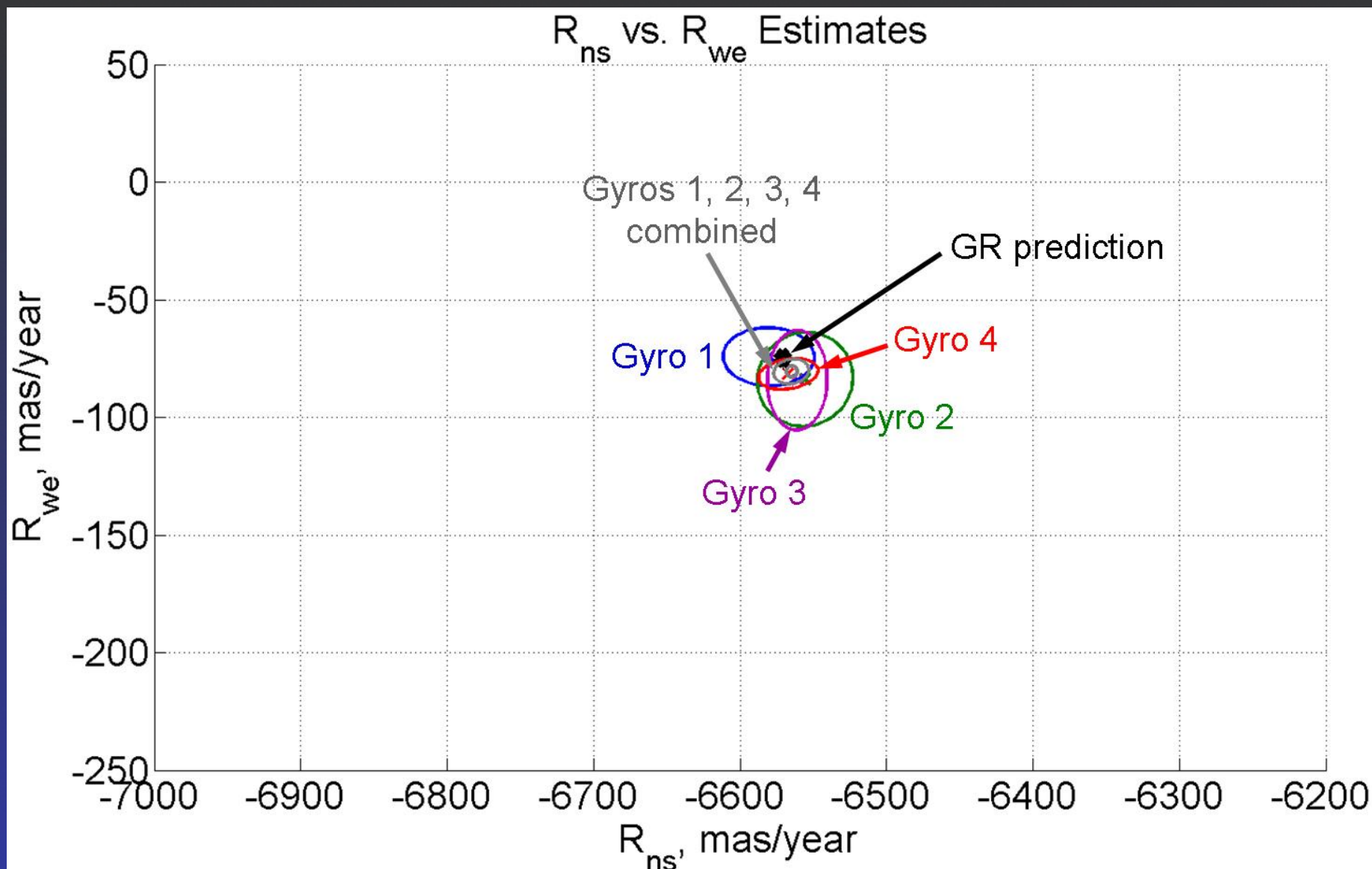


E-W (Frame-dragging)

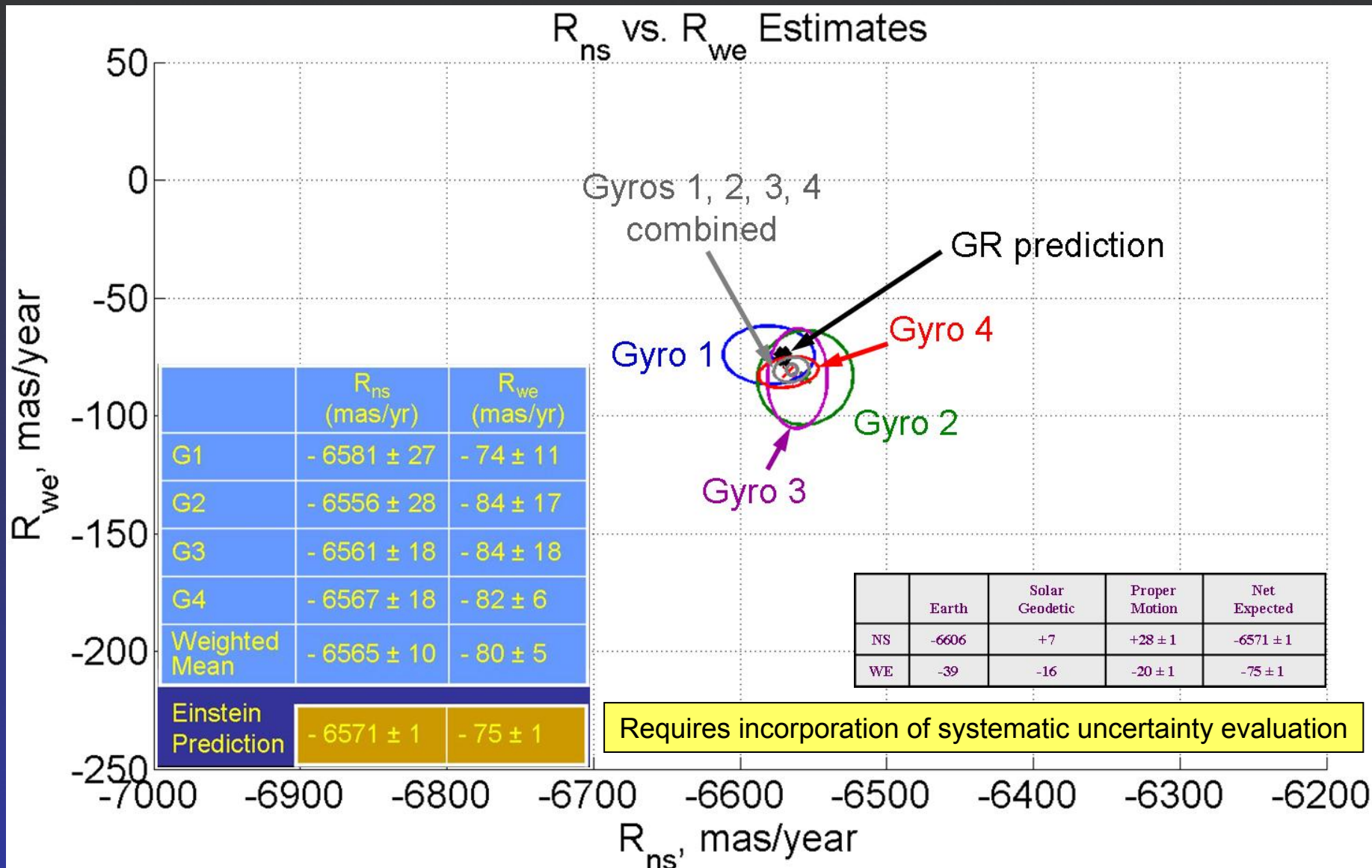


(note: different y-axis scale for N-S vs. E-W)

Full Model Results as of Dec '08



Full Model Results as of Dec '08 - II



Independent SuperGeometric Results

1 R_{NS} & R_{WE} – Gyro #4

$$R_{NS} = -663 \pm 16 \text{ mas/yr}$$

$$R_{WE} = -77 \pm 14 \text{ mas/yr}$$

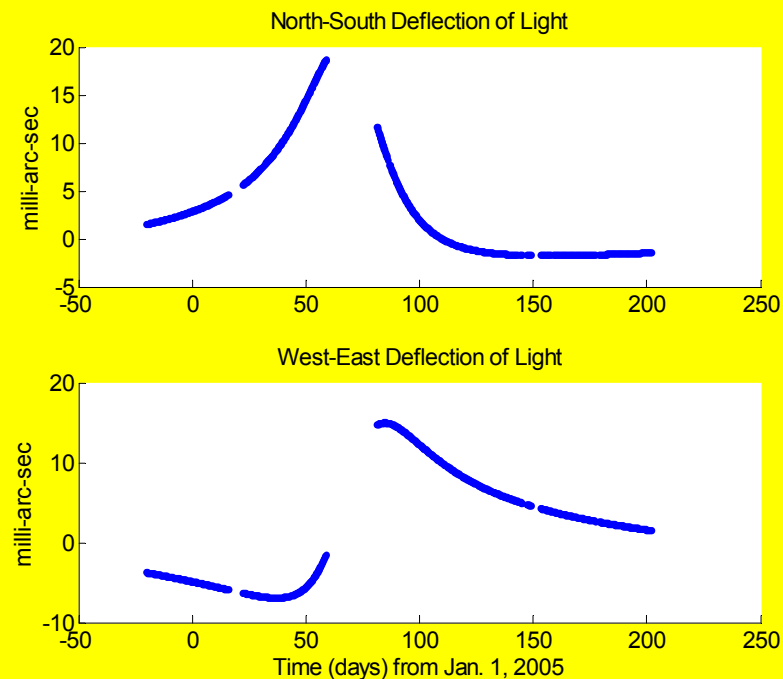
2 Gravitational deflection of light



3 24.6487-day guide star orbital motion

$$a_{NS} = -1.65 \pm 0.14 \text{ mas} \quad b_{NS} = -0.36 \pm 0.14 \text{ mas}$$

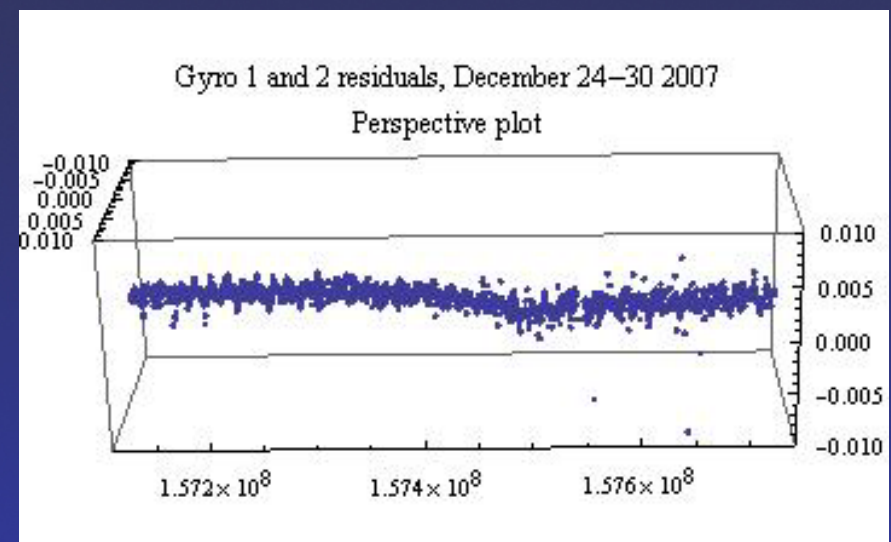
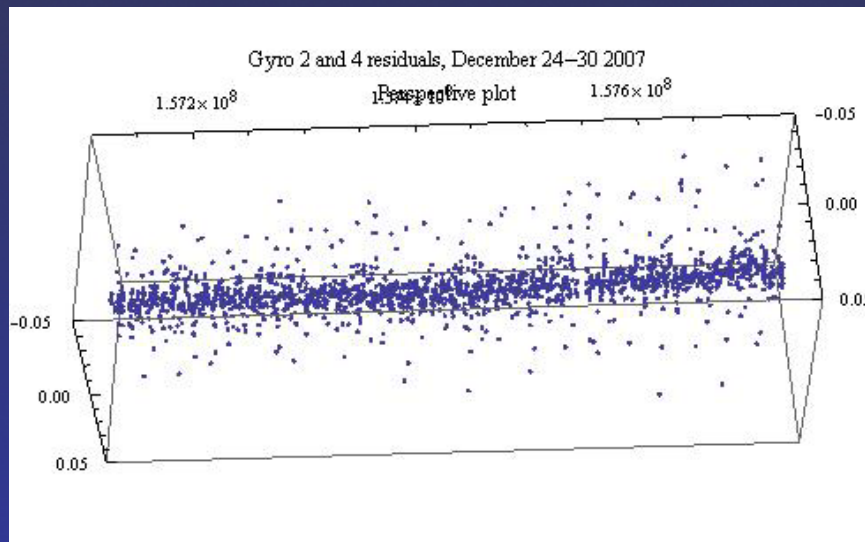
$$a_{WE} = -0.11 \pm 0.12 \text{ mas} \quad b_{WE} = -0.61 \pm 0.13 \text{ mas}$$



Result to date $\lambda = 1.20 \pm 0.17$

Advanced Investigations of Systematics

- Improved science modeling
 - Exact treatment of readout nonlinearities
 - Thermal correlations
 - 2-sec processing
- Detailed gyro-to-gyro comparisons



Are we done?

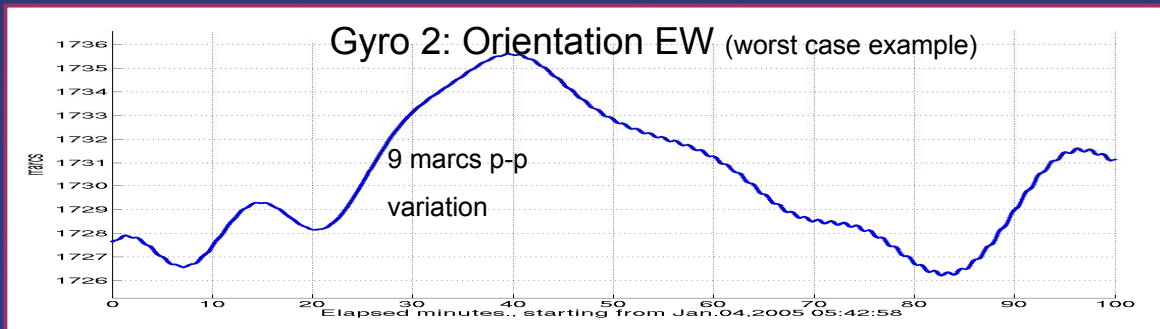
- **Current '2 floor' 4-gyro result ~ 5 marcs/yr (statistical uncertainty)**
 - limited by \longrightarrow $\left\{ \begin{array}{l} \text{once-per-orbit time step} \\ 3 \text{ out of 10 segments (158 days vs. 333 days)} \end{array} \right.$
- **Fundamental & operational limits**
 - SQUID noise 0.14 – 0.35 marcs/yr (gyro dependent)
 - Covariance ~ 1-2 marcs/yr (4 gyros combined)

$$\frac{ds_{NS}}{dt} = r_{NS} + k(t)\mu_{EW} + \sum A_m(t) \cos \Delta\phi_m - B_m(t) \sin \Delta\phi_m$$

$$\frac{ds_{EW}}{dt} = r_{EW} - k(t)\mu_{NS} + \sum A_m(t) \sin \Delta\phi_m + B_m(t) \cos \Delta\phi_m$$

$$\Delta\phi_m = \phi_{\text{roll}} - m\phi_p$$

Note: noiseless trajectory shown. \longrightarrow
Noise must be added for complete evaluation



Accurate integration requires time step \ll 1 orbit \longrightarrow **parallel processing**

Advanced (2-sec) Filter Development

Two simultaneous development activities

Co-PI Charbel Farhat



- 1 **Serial processing** → **accuracy**
 - ◆ built on 4-years' experience with 2-floor (once per orbit) filter
 - ◆ accurately models the physical phenomena
 - ◆ has passed truth test
- 2 **Parallel processing** → **speed**
 - ◆ ~ 10x faster (3 day analysis → overnight)
 - ◆ requires computer cluster
 - ◆ ~ 10% modified/additional code

Talk by Michael Heifetz July 29



Michael Heifetz



Badr Alsuwaidan



Majid Almeshari



John Conklin



Vladimir Solomonik

Upcoming GPB Talks

- June 24
Sasha Buchman
Mac Keiser
Evidence for patch effect
Torques & analysis treatment
- July 1
Alex Silbergleit
Trapped flux mapping
- July 29
Michael Heifetz
Data Analysis Past, Journey

Path to Completion

- Serial 2-sec processing (applied to data subset) *Aug '09*
- Complete transition to parallel processing *Oct '09*
- Extension from 5, 6, 9 (applied to full data set) *Dec '09*
- Complete treatment of systematics *Jan '10*
- Blind test against SAO guide star orbital motion *Feb '10*
- Grand synthesis of Geometric & Algebraic results
 ➡ approach ~ 1-2 marcs/yr 4-gyro limit *Jun '10*

**Final results to be announced at Fairbank Workshop on
Fundamental Physics & Innovative Engineering in Space**

Aug '10